Model STV Digital Integrating Video Camera and Autoguider

The *STV* is an entirely new concept for video imaging and autoguiding. The camera is a hybrid design that combines the best of both traditional video technology and cooled integrating digital CCD camera technology. The result is a cooled, integrating digital video camera capable of recording much fainter objects than traditional video cameras and at the same time give the performance of a cooled integrating CCD camera in a self-contained unit that is as easy to use as a video camera at the telescope.

SBIG WARKEN
STV Digital Video Camera and Autoguider

Modern astronomical CCD cameras use cooled detectors to lower the dark current and the electronics are designed to integrate (take exposures) up to an hour long for a single exposure. This technology results in very sensitive cameras capable of reaching greater than 20th magnitude with typical amateur telescopes from backyard sites. Our web site is full of beautiful images from amateurs all over the world who have captured many dim deep space objects with amazing detail. But CCD imaging is not as simple as using a video camera. One must use a computer of some kind at the telescope to control the CCD camera and save the images. This extra bit of equipment requirement and additional power cords, etc., is often just enough to make CCD imaging too much of a hassle for some. However, with the *STV*, this is also no longer a problem. No computer is required to use the STV as an autoguider or an imager.

What is the STV?



M27. RGB color image taken with STV equipped with CFW5C internal color filter wheel. *Alan Holmes / SBIG*

Video astronomy is a popular pursuit because nearly everyone is familiar with video cameras. The idea of putting a video camera at the eyepiece of a telescope and seeing the results on a TV monitor is very exciting. It is an easy way to share the views of the night sky with friends and family. The viewing is comfortable and the images can be recorded on tape. But one limitation of typical video astronomy is the inability of even "low light" video cameras to capture faint nebula, galaxies or stars dimmer than about 9th or 10th magnitude through an 8" scope. By cooling the detectors and designing the cameras to take long exposures (integrate) over many minutes, modern astronomical CCD cameras like the ST-7 have overcome much of the limitation of video astronomy. However, video cameras remain an interesting accessory for amateur astronomers because of their relative ease of use, despite their inability to record dim objects.

This limitation of typical video cameras is a result of the common video format. Video cameras cannot usually take exposures longer than 1/30th of a second. Even if they could, the CCD detectors used in typical video cameras are not cooled. The dark current that would build up in longer exposures would result in noisy looking images. The only way to increase the sensitivity

of typical video cameras up to now has been to add an image intensifier (e.g., night vision scope) between the camera and the telescope. This works, to a degree. But the images are usually very grainy and suffer from poor resolution. The result is that amateur video astronomy is generally limited to imaging bright objects like the moon and planets. This problem has been solved in the *STV*.



By utilizing a cooled detector and designing all new electronics around a DSP (digital signal processor) we have been able to achieve the performance of a cooled integrating astronomical CCD camera in a more traditional video format. Exposures from 1/1000th of a second to 600 seconds are possible. The video output can be viewed on the built-in LCD display or on any external video monitor. We estimate the STV will reach 14th magnitude stars with one second integrations through an 8" SCT - and 18th magnitude in 60 seconds. Objects that are impossible to capture with a typical low light video camera are easily targets for the *STV*. Globular clusters, nebula and galaxies are easily displayed on a video monitor, recorded on video tape or captured as digital image files. One simply increases the exposure until the objects are visible and presses a button to capture the image.

How does it work?

The **STV** operates very much like a traditional video camera but with some important distinctions. The **STV** consists of a camera head and a control box. The camera head contains the detector, some electronics and the thermoelectric cooler. The head is connected to the control box by a 15 foot cable. The control box requires 12VDC. The camera head has a 1.25" nose piece which means it will fit any telescope or finder scope that accepts.



1.25" eyepieces. It can be used in place of an eyepiece in the main telescope, or it can be used as an "electronic" eyepiece in a finder scope. When used as an autoguider, the **STV** head will fit in a separate guide scope or may be used in place of a cross-hair guiding eyepiece in most typical off-axis guider assemblies. And, since the **STV** is approximately 30 times as sensitive as an ST-4 autoguider, finding suitable guide stars in an off-axis guider through the main OTA will be a much easier task.



No external computer or video display is required to operate the **STV**, and when used as an autoguider, the optional built-in 5" LCD display is not needed although it may make finding and focusing stars much easier. The **STV** has a 2 line x 24 character alphanumeric display that provides all the information needed to operate as a stand alone autoguider. In this configuration it is operated similar to the older ST-4.

What does it do?

STV Image of the Horsehead Nebula taken with Celestron 8" Fastar. *Alan Holmes / SBIG*

The STV is a Digital Integrating Video Camera and Autoguider:

Digital Integrating Video: The **STV** will take digital video images and instantly display them on the built-in 5" LCD display

or on any external video monitor. "Integrating" means that the camera is capable of exposure times much longer than 1/30th of a second. The images can be recorded on video tape. The update rate for the video display depends on the mode selected and of course the length of the exposure. If one selects exposures longer than say one second, the display will not update until the exposure is completed. However, the previous frame remains on the video display until it is replaced by a new image. For short exposures (e.g., 1/1000th of a second) the display will update as fast as 16 frames per second.

Camera: In addition to digital video output, the **STV** will take and store digital images in its on-board memory just like a digital camera. Stored images may be downloaded to a computer at a later time. If one wishes to connect a computer to the **STV** while it is in operation, digital images may be downloaded to the computer with the push of a button on the **STV** control box. Once transferred to a computer, the images may be viewed, processed, etc., with CCDOPS software or exported to a popular image file format such

as TIFF for use with third party image processing software.

Autoguider: The **STV** will also autoguide! We now have 10 years of experience making the leading autoguider in the world for astrophotography and this experience has taught us a thing or two about how to make the task a little easier. All these improvements are incorporated in the **STV**. Believe it or not this was its *sole* intended purpose when it was first put on the drawing board two years ago. It has obviously evolved far beyond just an autoguider - but these functions are still there. As a stand alone autoguider (no computer required) the **STV** will reach several magnitudes fainter than the old "work horse" ST-4.



Electronic Finder: If you already have an astronomical CCD camera and have difficulty placing objects on the detector because they can't be seen, the STV will make a great accessory as an "electronic finder." If placed at the eyepiece tube of a finder scope the STV can record and display star fields with limiting magnitudes comparable to common finder charts or the exposure can be increased until the deep space object is actually seen through the finder scope. This can eliminate the need for removing the main CCD camera from the telescope to find and center objects on its detector.

Seeing Monitor:

The STV will generate a graph on the video display showing the seeing conditions as a function of the FWHM of star images as sampled every 2/10ths of a second.

Telescope Drive Monitor:

The STV will generate a graph on the video display showing the error in RA and DEC as your telescope drive operates in real time.

Display Modes:

A partial frame mode with either normal or zoom display may be selected for faster focusing. This mode results is update rates about 2X faster than normal imaging mode. The time and date are not displayed in focus or analysis mode.

In WIDE FIELD display mode, the full 656 x 480 pixel array is binned 3x3 and displayed so the entire field of view of the detector is visible on the video monitor. The time and date are displayed in the lower right hand corner of the video frame outside the image area.



In NORMAL display mode a 640×400 portion of the detector is binned 2x2 and displayed so that the full width of the image fills the width of the video monitor. The time and date are displayed in the lower right hand corner of the frame outside the image area.

In ZOOM display mode the central 320×200 portion of the detector is displayed unbinned so that the width of the image fills the width of the video monitor. The time and date are displayed in the lower right hand corner of the



frame outside the image area.

One unique imaging mode is the grid display. In this mode a video overlay of 6 lines yields 9 handy cross hairs that may be used for visual placement of guide stars when using the STV for manual guiding. The time and date are displayed in the lower right hand corner of the frame outside the image area.

In ANALYSIS mode, the alphanumeric display provides several types of information including the X,Y location of the moving cross-hair as well as the value of the pixel under the cross-hair and the average value. Large movable cross-hairs superimposed on the video monitor are controlled by two rotary dials on the STV. One places the intersection anywhere on the image and presses a button. The first selection measures the background and automatically marks the location

with a box. The large cross hairs are then moved to a star or other location and the digital display automatically gives the magnitude of any star placed under the cross hairs. If the select button is pressed again, this second location is marked by a small cross hair. When the large cross hairs are moved to a third location the digital readout on the STV automatically displays the separation in arc seconds from the previous star or location marked by the small cross-hairs.

In addition to the various display modes described above, the STV has several analytical modes which allow you to measure seeing, periodic error and guiding accuracy:

STVREMOTE

A PC may be used to remotely control the STV. To transfer images to a PC or to remotely control the STV a serial cable must be connected between the Serial I/O (RS232) port on the STV and COM1 or COM2 serial port on a PC. STVREMOTE software is included with every STV. This software creates a virtual STV on your computer monitor. By clicking on the corresponding buttons on the computer screen, virtually every function of the STV is duplicated and controlled from the user's keyboard.



Images that have been saved in the STV's internal memory are downloaded to the user's computer using STVREMOTE software. Once downloaded, the images may be displayed on the virtual STV's screen or saved in a variety of formats for viewing by other software programs like CCDOPS or CCDSOFTV5.

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OPTIONS

CFW5C Internal Color Filter Wheel

One recent improvement in the STV is the ability to control an internal color filter wheel. This allows the user to take tri-color or quad color images in similar fashion to other CCD cameras. The same high quality color filter wheel designed for the ST-237A CCD camera will fit in the STV head.

FR237 Focal Reducer and eFinder Assembly

Our eFincer is a highly recommended accessory. It consists of a focal reducer that doubles as a wide angle objective lens when used with the eFinder tube. The focal reducer lens will screw into the nosepiece of the STV to increase the field of view and decrease the effective f/ratio of your optical system. It is designed to be used on systems with focal ratios of F/15 to F/6.3. For example, when used with and without the extension tube the focal reducer will yield the following focal ratios:

Focal ratio of parent system	F/10	F/6.3
Resulting focal ratio without extension tube	F/5.95	F/3.75
Resulting focal ratio with extension tube	F/3.75	Not recommended

The same focal reducer lens is also designed to work as a convenient wide field finder lens when placed at the correct distance from the CCD. The eFinder assembly screws into the STV head in place of the 1.25" nosepiece and holds the focal reducer lens at the correct distance. In this configuration the STV becomes a very sensitive "electronic finder" with a field of view of approximately 2.7 degrees. Any star in the Uranometria catalog can be seen in a three second exposure through the eFinder lens. In addition, the STV will autoguide with this lens assembly to an accuracy of *less than one arcsecond*, eliminating the need for a separate guide scope.



This is the first cooled integrating video CCD camera and autoguider offered to the amateur astronomy market. The **Standard STV** includes the thermoelectrically cooled camera head with standard T-threads, removable 1.25" nose piece, control box with alphanumeric display, 110VAC adapter, and internal neutral density filter. The **Deluxe STV** includes the all the standard items plus a built-in 5" LCD video display, tripod mounting foot and a custom hard carrying case.

STV Typical Specifications		
 Complete Stand Alone Operation Video Imager with on board video display or external video monitor Autoguider (30 times more sensitive than an ST-4) Digital Imager (using on board memory for image storage) 	 Self-contained Control Panel: 11 button keyboard with two rotary switches for setting numbers 2 line x 24 character vacuum fluorescent digital display 5 inch diagonal LCD TV screen built in (optional) 	
 Video or Digital Image Formats Video output for connection to any TV or VCR RS232 output for downloading digital images to any computer 	Focus Mode: - sub-frame and electronic zoom features	
Image Mode: - exposure times from 0.001 to 600 seconds - full or zoom format - single or continuous update - Track and Accumulate (Patented)	Moni <i>tor Mode</i> : - measure atmospheric seeing - measure telescope drive periodic error	
Display/Crosshairs Mode: - adjust display brightness and contrast - measure star magnitudes - measure stellar separations	Track Mode: - automatic or user select - guide to: current position, last position, displayed cursor or selected star - aggressiveness 0.1 to 2.0	
File I/O Mode: - display image - save file to memory - download displayed image - download all images in memory - erase displayed image - erase all images in memory	Calibrate Mode: - automatic or user select - exposure times 0.01 to 99.9 seconds - movement: incremental or single - calibration time: 0.1 to 9.99 - active relays: X & Y, X only, Y only - view calibration results	
Set Up Mode: - set time / date - show grid - night vision (on/off) - video display (yes/no) - chose units of measure (in/cm) - set focal length - set aperture - telescope type (refractor or reflector) - set site ID (1 to 155) - show video test pattern 1 (box to show size) - show video test pattern 2 (gray scale, 0 to max) - set video offset for dark level	Option Mode: - self test - open shutter - close shutter - run shutter - show CCD temp - measure image (average and rms noise)	
Video Output: - NTSC Standard - 525 horizontal lines with 720 pixels / line - PAL Standard available as a menu option	CCD: - Texas Instruments TC-237 - High Sensitivity Frame Transfer (electronic shutter) CCD - 656 x 480 pixels @ 7.4 microns square	
Image Modes: - Normal = 640 x 480 binned 2x2 - Wide = 656 x 480 binned 3x3 - Zoom = 320 x 200 binned 1x1	Image bit depth: - 10 bits for 1x1 binning mode - Up to 16 bits for other binning modes and Track and Accumulate images	
CPU Weight: 4lb (1.8kg) with LCD, 3.2lb (1.4kg) without LCD	Head Weight: 16oz (0.4kg)	
CPU Dimensions: 11.6 x 9.4 x 2.6 inches (29.5 x 23.7 x 6.7 cm	Head Dimensions: 3.25 x 1.75 inches (8.3 x 4.4 cm)	

Specifications are subject to change without notice